

Construction of hyperboloidal initial data without logarithmic singularities

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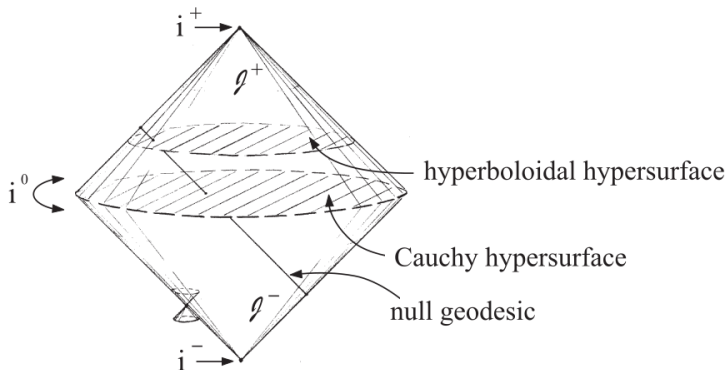
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Penrose 1963: asymptotically simple spacetimes & conformal compactification

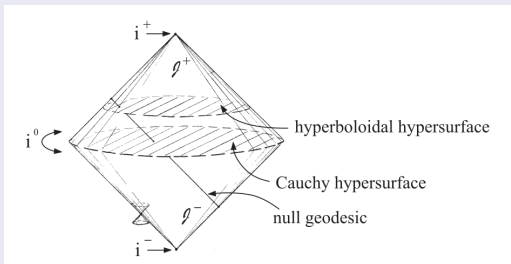


the asymptotic behavior of the gravitational field of an isolated system resembles that of Minkowski space

- There exists a **smooth** manifold $\widetilde{M} = M \cup \partial M$ with boundary,
- and a **smooth** Lorentz metric $\widetilde{g} = \Omega^2 g$ on \widetilde{M}
- such that $\Omega = 0$, $d\Omega \neq 0$ on ∂M
 - The **degree of smoothness** of the involved structures **is critical**.

Asymptotically simple spacetimes & asymptotically hyperboloidal data

- **Friedrich:** Solutions to the **hyperboloidal initial value formulation of his conformal field equations** are asymptotically simple in the future.
 - Hyperboloidal initial data – sufficiently and close to Minkowskian hyperboloidal data – develop into solutions which admit smooth conformal extensions containing a regular point i^+ that represents future time-like infinity.



the asymptotic behavior of the gravitational field of an isolated system resembles that of Minkowski space [Credit: H.Friedrich]

- An **asymptotically hyperboloidal vacuum** data set (Σ, h_{ij}, K_{ij}) is such that
 - Σ is the interior of a **compact manifold with boundary** $\Sigma \cup \partial\Sigma$
 - the **trace** $K = K_{ij}h^{ij}$ of K_{ij} is **bounded away from zero** near $\partial\Sigma$
 - if ω is a defining function for $\partial\Sigma$ then $\omega^2 h_{ij}$ and $\omega \mathring{K}_{ij}$ **extend regularly** to $\partial\Sigma$

Constraint equations

Einstein equations: $E_{ab} = G_{ab} - \mathcal{G}_{ab} = 0$

$$n^a n^b E_{ab} = 0 \text{ \& } \pi_i^a n^b E_{ab} = 0$$

- in the vacuum case ($\mathcal{G}_{ab} = 0$) for (h_{ij}, K_{ij}) on Σ

$$^{(3)}R + (K_{ij}h^{ij})^2 - K_{ij}K^{ij} = 0 \text{ \& } D^i [K_{ij} - h_{ij} (K_{ef}h^{ef})] = 0 \quad (D_k h_{ij} = 0)$$

- It is an underdetermined system, 4 equations for the 12 variables: (h_{ij}, K_{ij})

Conformal method (Lichnerowicz 1944, York 1972)

$$h_{ij} = \phi^4 \tilde{h}_{ij}, \quad \overset{\circ}{K}_{ij} = K_{ij} - \frac{1}{3} h_{ij} K = \phi^{-2} \tilde{K}_{ij}, \quad \tilde{K}_{ij} = CKO[X]_{ij} + \tilde{K}_{ij}^{[TT]}$$

$$(h_{ij}, K_{ij}) \longleftrightarrow (\phi, \tilde{h}_{ij}; K, \mathbf{X}_i, \tilde{K}_{ij}^{[TT]})$$

the constraints form a quasilinear elliptic system for (ϕ, \mathbf{X}_i)

Evolutionary method (I.R. 2015)

Σ is foliated by topological two-spheres & applying an ADM type decomposition

$$h_{ij} \longleftrightarrow \hat{N}, \hat{N}^i, \hat{\gamma}_{ij} \quad \& \quad K_{ij} \longleftrightarrow \kappa, \mathbf{k}_i, \mathbf{K}_{ij} [= \overset{\circ}{\mathbf{K}}_{ij} + \frac{1}{2} \mathbf{K} \hat{\gamma}_{ij}]$$

using a complex null dyad q^i : $q_{ij} = q_{(i} \bar{q}_{j)}$ $(h_{ij}; K_{ij}) \longleftrightarrow (\hat{\mathbf{N}}, \mathbf{N}, \mathbf{a}, \mathbf{b}; \kappa, \mathbf{k}, \mathbf{K}, \overset{\circ}{\mathbf{K}}_{qq})$

the constraints form a parabolic-hyperbolic system for $(\hat{\mathbf{N}}, \mathbf{k}, \mathbf{K})$

Andersson & Chruściel '93, '94, '96: log-terms become part of the discussions

- They proved that **even if the free data is smooth** on $\tilde{\Sigma} = \Sigma \cup \partial\Sigma$,

- (1) **in general**, the constrained fields have poly-logarithmic expansions in $\omega \sim \rho^{-1}$, where ρ denotes the “distance” from the isolated system

$$\mathcal{C} = \mathcal{C}_0 + \mathcal{C}_1\omega + \mathcal{C}_2\omega^2 + \dots \quad \mathcal{C} = \mathcal{C}_0 + \sum_{i=1}^{\infty} \omega^i \left[\mathcal{C}_i + \sum_{j=1}^{\mathcal{N}_j} \mathcal{C}_{i,j}^{[log]} \log^j \omega \right]$$

$$(n \geq 1) \quad \omega \log^n \omega \rightarrow 0 \quad \text{but} \quad \partial_\omega(\omega \log^n \omega) \text{ blows up}$$

- (2) **non-generic cases**: the initial data can be smooth (free of log-terms) on $\tilde{\Sigma}$ if the relations

$$\tilde{K}_{ra}^{[log]}|_{\partial\Sigma} = 0 \quad \text{in} \quad \tilde{K}_{ab} = \tilde{K}_{ab}^{[C^\infty]} + \Omega^2 \log \Omega \cdot \tilde{K}_{ab}^{[log]} \quad \& \quad (\tilde{K}_{ab}^\circ - \tilde{\tilde{K}}_{ab}^\circ)|_{\partial\Sigma} = 0$$

(derived from the free data) **hold on** $\partial\Sigma$.

Facing the problems:

- These conclusions were disappointing because they implied that the initial data constructed by the conformal method is **generally not regular enough** for use in Friedrich's existence theorems. $[\tilde{\nabla}_e d_{abc}{}^e = 0 \quad \& \quad d_{abc}{}^e = \Omega^{-1} C_{abc}{}^e]$
- **If the initial data involves log-terms, then the development will also contain**
- How can then the metric be decomposed into a sufficiently smooth non-physical metric and a conformal factor? the concept of asymptotically simple spacetimes?

Important new results by Beyer and Ritchie [CQG,39,145012 (2022)]

- They assumed that there exist smooth global solutions to the **parabolic-hyperbolic form of the constraints** on a "hyperboloidal initial data surface" Σ .
- They also assumed that these solutions **extend regularly up to order 4 and 3**, respectively, to $\partial\Sigma$
 - $\hat{\mathbf{N}}, \mathbf{K} \in C^4([0, \omega_0), C^\infty(\mathcal{S}^2))$
 - $\mathbf{k} \in C^3([0, \omega_0), C^\infty(\mathcal{S}^2))$
- Then, using an **impressive Fuchsian-equation based argument** they showed that **the constrained variables extend smoothly** to $\partial\Sigma$, whence such solutions are free of all log-terms.
- **parabolic-hyperbolic form of the constraints** ✓
 - BUT **even stronger** assumptions than used by Andersson and Chruściel.
 - They restricted not only the free data (that went largely uncommented)
 - **but** also **the constrained fields** $\hat{\mathbf{N}}, \mathbf{K}, \mathbf{k}$!!!

The strategy we used:

Asymptotically hyperboloidal data: $(h_{ij}; K_{ij}) \longleftrightarrow (\hat{\mathbf{N}}, \mathbf{N}, \mathbf{a}, \mathbf{b}; \boldsymbol{\kappa}, \mathbf{k}, \mathbf{K}, \mathring{\mathbf{K}}_{qq})$

- Recall first that a data set $(\hat{\mathbf{N}}, \mathbf{N}, \mathbf{a}, \mathbf{b}; \boldsymbol{\kappa}, \mathbf{k}, \mathbf{K}, \mathring{\mathbf{K}}_{qq})$ can only be **asymptotically hyperboloidal** if the following falloff conditions hold: $\omega \sim \rho^{-1}$

$$\begin{aligned} \hat{\mathbf{N}} &= \hat{\mathbf{N}}_1 \omega + \mathcal{O}(\omega^2) & \mathbf{K} - 2\boldsymbol{\kappa} &= \mathcal{O}(\omega) & \mathbf{k} &= \mathcal{O}(1) \\ \mathbf{a} &= \omega^{-2} + \mathcal{O}(\omega^{-1}) & \mathbf{b} &= \mathcal{O}(\omega^{-1}) & \mathbf{N} &= \mathcal{O}(\omega) & \mathring{\mathbf{K}}_{qq} &= \mathcal{O}(\omega^{-1}) \end{aligned}$$

The free data is assumed to be smooth on $\Sigma \cup \partial\Sigma$: $C^\infty([0, \omega_0), C^\infty(\mathbb{S}^2))$

$$\begin{aligned} \mathbf{N} &= \mathbf{N}_1 \omega + \mathbf{N}_2 \omega^2 + \mathcal{O}(\omega^3) \\ \mathbf{a} &= \omega^{-2} + \mathbf{a}_{(-1)} \omega^{-1} + \mathbf{a}_0 + \mathbf{a}_1 \omega + \mathbf{a}_2 \omega^2 + \mathcal{O}(\omega^3) \\ \mathbf{b} &= \mathbf{b}_{(-1)} \omega^{-1} + \mathbf{b}_0 + \mathbf{b}_1 \omega + \mathbf{b}_2 \omega^2 + \mathcal{O}(\omega^3) \\ \boldsymbol{\kappa} &= \boldsymbol{\kappa}_0 + \boldsymbol{\kappa}_1 \omega + \boldsymbol{\kappa}_2 \omega^2 + \mathcal{O}(\omega^3) \\ \mathring{\mathbf{K}}_{qq} &= \mathring{\mathbf{K}}_{qq(-1)} \omega^{-1} + \mathring{\mathbf{K}}_{qq0} + \mathring{\mathbf{K}}_{qq1} \omega + \mathring{\mathbf{K}}_{qq2} \omega^2 + \mathcal{O}(\omega^3) \end{aligned}$$

Use the most generic poly-logarithmic form of the constrained fields $(\hat{\mathbf{N}}, \mathbf{K}, \mathbf{k})$:

$$\begin{aligned} \hat{\mathbf{N}} &= \sum_{i=1}^{\infty} \omega^i \left[\hat{\mathbf{N}}_i + \sum_{j=1}^{\mathcal{N}_j} \hat{\mathbf{N}}_{i,j}^{[log]} \log^j \omega \right], & \mathbf{K} &= \mathbf{K}_0 + \sum_{i=1}^{\infty} \omega^i \left[\mathbf{K}_i + \sum_{j=1}^{\mathcal{N}_j} \mathbf{K}_{i,j}^{[log]} \log^j \omega \right] \\ \mathbf{k} &= \mathbf{k}_0 + \sum_{i=1}^{\infty} \omega^i \left[\mathbf{k}_i + \sum_{j=1}^{\mathcal{N}_j} \mathbf{k}_{i,j}^{[log]} \log^j \omega \right], \text{ where } \hat{\mathbf{N}}_1 = \boldsymbol{\kappa}_0^{-1}, \mathbf{K}_0 = 2\boldsymbol{\kappa}_0, \mathbf{k}_0 = \boldsymbol{\kappa}_0^{-1} \delta \boldsymbol{\kappa}_0 \end{aligned}$$

Our first main result

Theorem I:

- Choose **free data** $(N, a, b, \kappa, \overset{\circ}{K}_{qq})$ on Σ such that it **satisfies the asymptotically hyperboloidal falloff conditions** and **it is smooth on $\Sigma \cup \partial\Sigma$** .
- Suppose that (\hat{N}, K, k) are **smooth solutions** of the parabolic-hyperbolic form of the constraints on Σ .
- (\hat{N}, K, k) are also assumed to possess the most generic poly-logarithmic expansion near $\partial\Sigma$ as indicated above.
- Then the corresponding asymptotically hyperboloidal initial data set **admits well-defined Bondi energy and angular momentum if and only if** all coefficients of the **logarithmic terms vanish up to order four and three** for \hat{N}, K and k , respectively, and, **in addition,**

$$\overset{\circ}{K}_{qq(-1)} = 0, \quad b_{(-1)} = 0, \quad \kappa_1 = 0.$$

The finiteness of the Bondi energy:

The Bondi energy can be given as the $\rho \rightarrow \infty$ limit of the Hawking (mass) energy

$$m_H = \sqrt{\frac{\mathcal{A}}{16\pi}} \left(1 + \frac{1}{16\pi} \int_{\mathcal{S}_\rho} \underbrace{(\mathbf{K}^2 - \hat{\mathbf{K}}^2 \hat{\mathbf{N}}^{-2})}_{=\Theta^{(+)}\Theta^{(-)}} \sqrt{\mathbf{d}} \overset{\circ}{\epsilon} \right)$$

$\mathcal{A}, \sqrt{\mathbf{d}} \sim \rho^2, \hat{\mathbf{K}} = \partial_\rho \log[\sqrt{\mathbf{d}}] \sim \rho^{-1}$ **requires** $\int_{\mathcal{S}_\rho} (\mathbf{K}^2 - \hat{\mathbf{K}}^2 \hat{\mathbf{N}}^{-2}) \sqrt{\mathbf{d}} \overset{\circ}{\epsilon} = -16\pi + \mathcal{O}(\rho^{-1})$

$\lim_{\rho \rightarrow \infty} m_H$ is finite \implies for all $j = 1, 2, \dots, \mathcal{N}_j$

$$\hat{\mathbf{N}}_{1,j}^{[log]} = \hat{\mathbf{N}}_{2,j}^{[log]} = \mathbf{K}_{1,j}^{[log]} = 0$$

$$\mathbf{K}_{2,i}^{[log]} = \mathbf{K}_0 (2 \mathbf{K}_{3,i}^{[log]} + \mathbf{K}_0^2 \hat{\mathbf{N}}_{4,i}^{[log]}) \cdot [\mathbf{a}_{(-1)} \mathbf{K}_0 + 4 \mathbf{K}_1]^{-1}$$

$$\hat{\mathbf{N}}_{3,i}^{[log]} = -2 (2 \mathbf{K}_{3,i}^{[log]} + \mathbf{K}_0^2 \hat{\mathbf{N}}_{4,i}^{[log]}) \cdot (\mathbf{K}_0 [\mathbf{a}_{(-1)} \mathbf{K}_0 + 4 \mathbf{K}_1])^{-1}$$

The finiteness of the Bondi angular momentum: I.R.

- details of the construction in arXiv:2401.14251 [to appear in PRD]
- for any axial vector field ϕ^a

$$J[\phi] = -(8\pi)^{-1} \int_{\mathcal{S}_\rho} (\phi^a [\bar{q}_a \mathbf{k} + q_a \bar{\mathbf{k}}]) \sqrt{\mathbf{d}} \overset{\circ}{\epsilon}$$

- since $\sqrt{\mathbf{d}} \sim \rho^2$ the Bondi angular momentum cannot be finite unless for all $j = 1, 2, \dots, \mathcal{N}_j$

$$\mathbf{k}_{1,j}^{[log]} = \mathbf{k}_{2,j}^{[log]} = 0$$

Completion of the proof

- To obtain the desired restrictions, we substitute the updated form of the asymptotic expansions into the **parabolic-hyperbolic system** and sort the terms with respect to powers of ρ^{-1} and also of $\log \rho$. \implies

$$\mathbf{k}_{1,j}^{[log]} = \mathbf{k}_{2,j}^{[log]} = \mathbf{k}_{3,j}^{[log]} = 0, \quad \begin{aligned} \hat{\mathbf{N}}_{1,j}^{[log]} &= \hat{\mathbf{N}}_{2,j}^{[log]} = \hat{\mathbf{N}}_{3,j}^{[log]} = \hat{\mathbf{N}}_{4,j}^{[log]} = 0 \\ \mathbf{K}_{1,j}^{[log]} &= \mathbf{K}_{2,j}^{[log]} = \mathbf{K}_{3,j}^{[log]} = \mathbf{K}_{4,j}^{[log]} = 0 \end{aligned}$$

Inspecting the consequences

- The above findings suggest using the following **asymptotic expansions** in the parabolic-hyperbolic form of constraints **for constrained variables**:
 - $\hat{\mathbf{N}} \longrightarrow \hat{\mathbf{N}}_0 + \hat{\mathbf{N}}_1 \omega + \hat{\mathbf{N}}_2 \omega^2 + \hat{\mathbf{N}}_3 \omega^3 + \hat{\mathbf{N}}_4 \omega^4 + \omega^4 w_{\hat{\mathbf{N}}}(\omega)$
 - $\mathbf{K} \longrightarrow \mathbf{K}_0 + \mathbf{K}_1 \omega + \mathbf{K}_2 \omega^2 + \mathbf{K}_3 \omega^3 + \mathbf{K}_4 \omega^4 + \omega^4 w_{\mathbf{K}}(\omega)$
 - $\mathbf{k} \longrightarrow \mathbf{k}_0 + \mathbf{k}_1 \omega + \mathbf{k}_2 \omega^2 + \mathbf{k}_3 \omega^3 + \omega^3 w_{\mathbf{k}}(\omega)$
- The functions $w_{\hat{\mathbf{N}}}(\omega), w_{\mathbf{K}}(\omega), w_{\mathbf{k}}(\omega)$, are of class $C^0([0, \omega_0), C^\infty(\mathbb{S}^2))$ and **vanish at $\partial\Sigma$** , thus they can represent **all the higher-order log-terms** that may still occur.
- The coefficients $(\hat{\mathbf{N}}_4, \mathbf{k}_2, \mathbf{K}_1)$ represent the **asymptotic degrees of freedom** since
 - All of the “**bold-faced coefficients in black**” can be derived from the free data, $(\mathbf{N}, \mathbf{a}, \mathbf{b}; \kappa, \overset{\circ}{\mathbf{K}}_{qq})$, and from the asymptotic degrees of freedom $(\hat{\mathbf{N}}_4, \mathbf{k}_2, \mathbf{K}_1)$.

The asymptotic degrees of freedom and the asymptotic charges

- The terminology **asymptotic degrees of freedom** gets completely adequate by noticing that these quantities, $(\hat{\mathbf{N}}_4, \mathbf{k}_2, \mathbf{K}_1)$, together with the **free data**, $(\mathbf{N}, \mathbf{a}, \mathbf{b}; \overset{\circ}{\kappa}, \overset{\circ}{\mathbf{K}}_{qq})$, determine:

- The **Bondi energy**

$$m_B = \frac{1}{32\pi} \int_{\partial\Sigma} \left[\frac{\mathbf{K}_1^3}{\overset{\circ}{\kappa}_0} + 8 \hat{\mathbf{N}}_4 \overset{\circ}{\kappa}_0^3 - 4 \overset{\circ}{\kappa}_0 \overset{\circ}{\kappa}_3 + 12 \mathbf{a}_1 \overset{\circ}{\kappa}_0^2 - 2 \mathbf{K}_1 [\overset{\circ}{\kappa}_0^{-1} + 3 \mathbf{a}_0 \overset{\circ}{\kappa}_0 - 4 \overset{\circ}{\kappa}_2] \right. \\ \left. - \left\{ \overset{\circ}{\kappa}_0^5 \mathbf{N}_1^{-1/2} \bar{\partial} \left(\mathbf{N}_1^{3/2} \overset{\circ}{\kappa}_0^{-4} \mathbf{K}_1 \right) + 2 \overset{\circ}{\kappa}_0 \mathbf{N}_2^2 \bar{\partial} \left(\mathbf{N}_2^{-1} \overset{\circ}{\kappa}_0 \right) + \frac{1}{4} \overset{\circ}{\kappa}_0^5 \bar{\partial} \bar{\partial} \left(\overset{\circ}{\kappa}_0^{-6} \mathbf{K}_1 \right) \right. \right. \\ \left. \left. - \frac{3}{2} \overset{\circ}{\kappa}_0^{-2} \mathbf{K}_1 \bar{\partial} \bar{\partial} \overset{\circ}{\kappa}_0 + \frac{1}{4} \overset{\circ}{\kappa}_0 \bar{\partial} \bar{\partial} \mathbf{K}_1 + \text{“cc”} \right\} \right] \overset{\circ}{\epsilon}$$

- The components of the **Bondi angular momentum**

$$J_B[\phi_{(i)}] = -(16\pi)^{-1} \int_{\partial\Sigma} [(\phi_{(i)}^a \bar{q}_a) \mathbf{k}_2 + (\phi_{(i)}^a q_a) \bar{\mathbf{k}}_2] \overset{\circ}{\epsilon}$$

- where $\phi_{(i)}^a$, $(i = 1, 2, 3)$, denote the **axial Killing vector fields**

$$\overset{\circ}{\phi}_1^a = -\sin \varphi (\partial_\vartheta)^a - \cot \vartheta \cos \varphi (\partial_\varphi)^a \quad \& \quad \overset{\circ}{\phi}_2^a = \cos \varphi (\partial_\vartheta)^a - \cot \vartheta \sin \varphi (\partial_\varphi)^a \quad \& \quad \overset{\circ}{\phi}_3^a = (\partial_\varphi)^a$$

of a **centre-of-mass unit sphere reference system** $[\int_{\partial\Sigma} \vec{x} \overset{\circ}{\epsilon} = 0]$.

- The **algebraic conditions** on $\partial\mathcal{S}$: $\mathbf{a}_{(-1)} = \text{const} \ \& \ \overset{\circ}{\mathbf{K}}_{qq0} = \frac{1}{2} \kappa_0 \cdot \bar{\partial}\bar{\partial} \kappa_0^{-2} \ \&$
 - $\hat{\mathbf{N}} \longrightarrow \hat{\mathbf{N}}_0 + \hat{\mathbf{N}}_1 \omega + \hat{\mathbf{N}}_2 \omega^2 + \hat{\mathbf{N}}_3 \omega^3 + \hat{\mathbf{N}}_4 \omega^4 + \omega^4 w_{\hat{\mathbf{N}}}(\omega)$
 - $\mathbf{K} \longrightarrow \mathbf{K}_0 + \mathbf{K}_1 \omega + \mathbf{K}_2 \omega^2 + \mathbf{K}_3 \omega^3 + \mathbf{K}_4 \omega^4 + \omega^4 w_{\mathbf{K}}(\omega)$
 - $\mathbf{k} \longrightarrow \mathbf{k}_0 + \mathbf{k}_1 \omega + \mathbf{k}_2 \omega^2 + \mathbf{k}_3 \omega^3 + \omega^3 w_{\mathbf{k}}(\omega)$
- We get a **Fuchsian-type** (singular at $\omega = 0$) **equation** for the vector-valued variable $\underline{W} = (w_{\hat{\mathbf{N}}}, w_{\mathbf{K}}, w_{\mathbf{k}})^T$, comprised of the residuals, where $p \in \mathcal{S}^2$ and $0 < \omega < \omega_0$:

$$\partial_{\omega} \underline{W}(\omega, p) = \frac{1}{\omega} \text{diag}(0, -3, -1) \underline{W}(\omega, p) + \underline{H}\left(\omega, p; \hat{\mathbf{N}}_4(p), \mathbf{k}_2(p), \mathbf{K}_1(p), \underline{W}(\omega, p), \bar{\partial}\underline{W}, \bar{\partial}\bar{\partial}\underline{W}\right) \quad (*)$$

- where \underline{H} is a (lengthy, but **explicitly known**) vector-valued function that is **smooth in each of its arguments**, and **regularly extends to $\omega = 0$** .
- A “formal solution” of $(*)$ can be given by this integral equation

$$\underline{W}(\omega, p) = \text{diag}[1, \omega^{-3}, \omega^{-1}] \times \int_0^{\omega} \text{diag}[1, s^3, s] \times \underline{H}(s, p) \, ds \quad (**)$$

- Since the integrand regularly extends to $s = 0$, we can perform the change of variables in the integral by replacing s with the product $\omega \cdot \tau$, which yields

$$\frac{1}{\omega} \underline{W}(\omega, p) = \int_0^1 \text{diag}[1, \tau^3, \tau] \times \underline{H}(\omega \cdot \tau, p) \, d\tau \quad (***)$$

Our second main result:

Theorem II.

- Choose **free data** $(\mathbf{N}, \mathbf{a}, \mathbf{b}, \boldsymbol{\kappa}, \overset{\circ}{\mathbf{K}}_{qq})$ on Σ such that it **satisfies the asymptotically hyperboloidal falloff conditions** and it is **smooth** on $\Sigma \cup \partial\Sigma$.
- Suppose that $(\hat{\mathbf{N}}, \mathbf{K}, \mathbf{k})$ are **smooth** [i.e., of class $C^\infty((0, \omega_0), C^\infty(\mathbb{S}^2))$], solutions on Σ such that $\hat{\mathbf{N}} > 0$ there.
- Then, the constrained fields $(\hat{\mathbf{N}}, \mathbf{K}, \mathbf{k})$ are also of class $C^\infty([0, \omega_0), C^\infty(\mathbb{S}^2))$ on the whole of $\Sigma \cup \partial\Sigma$, i.e., **no logarithmic singularities occur, if and only if** the asymptotically hyperboloidal initial data set under consideration **admits well-defined Bondi energy and angular momentum**, and, **in addition**, for the free data

$$\mathbf{a}_{(-1)} = \text{const} \quad \& \quad \overset{\circ}{\mathbf{K}}_{qq0} = \frac{1}{2} \boldsymbol{\kappa}_0 \cdot \partial\partial \boldsymbol{\kappa}_0^{-2}$$

also hold on $\partial\Sigma$.

Summary:

- We proved that **the existence of well-defined Bondi energy and angular momentum**, together with some mild restrictions on the free data, implies that **the generic solutions of the parabolic-hyperbolic form of the constraint equations are free of logarithmic singularities**.
- We eliminated the restrictions imposed by Beyer and Ritchie on constrained fields. Additionally, we considerably weakened their restrictions on free data.
- Since Cauchy developments of smooth, asymptotically hyperboloidal initial data have smooth, conformal boundaries, **our result confirms** Penrose's **smoothness assumptions** used to define **asymptotically simple spacetimes**.
- Is there a way to **control the asymptotic charges** when constructing solutions to **the parabolic-hyperbolic form of the constraints**?
[Work in progress.]

Thanks for your attention